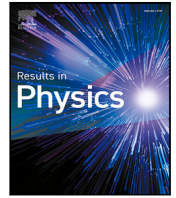


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Highly dispersive optical soliton perturbation with Kudryashov's sextic-power law nonlinear refractive index by semi-inverse variation

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ABSTRACT

A single bright highly dispersive optical soliton solution is recovered from semi-inverse variational principle. The model contains six power law nonlinear terms that constitute self-phase modulation effect which was proposed by Kudryashov. The amplitude–width relation of the soliton was finally recovered by Cardano's approach with appropriate and necessary constraints that are presented.

Introduction

A wide variety of features are studied in the context of optical solitons that travel down an optical fiber for inter-continental distances [1–20]. One of the most innovative concepts that has been conceived in telecommunications engineering is “highly dispersive” (HD) and “cubic–quartic” optical solitons. This was first reported in 2020 and 2021 respectively for preposterous consequences with low count of chromatic dispersion [3,12]. It was first studied for polarization-preserving fibers and later extended to birefringent fibers. For the scalar models, semi-inverse variational principle (SVP) was implemented to recover soliton solutions to a couple of nonlinear forms of refractive index. These are polynomial law and quadratic–cubic law. Today's paper moves a step further where HD solitons are studied with the newly proposed form of self-phase modulation (SPM), attributed to Nikolay Kudryashov. This form of SPM contains six nonlinear terms and is being referred to as sextic-power law nonlinearity [19]. The model

is studied with a few perturbation terms of Hamiltonian type, which show up maximum intensity.

Stepping back in time, it is worthwhile to mention that SVP was first proposed a couple of decades ago by J. H. He and all along this principle has gained its popularity and it has sustained ever since. This principle of SVP has been successfully applied to various other problems in optics, apart from Kudryashov's law of refractive index. Typically, it was studied in dispersive optical solitons, optical metamaterials, non-Kerr law solitons [11] including polynomial law and triple-power law [13]. It was also successfully applied to other scientific problems that had stemmed from Fluid Dynamics [7], Mathematical Biology [5] and Nuclear Physics [4]. It is not out of place to mention that this principle retrieved results in multi-dimensional optical solitons as well [9]. One of the shortcomings of this approach is that the retrieved soliton solution is not exact although it is analytical since this solutions is based on a certain principle. Although not an exact solution,

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it works when standard algorithms is rendered to be non-integrable for retrieving an exact solution. Subsequent sections illustrate the details.

Governing model

The governing equation having Kudryashov’s sextic form of non-linear refractive index with the inclusion of perturbation terms is structured as [19]:

$$i q_t + i a_1 q_x + a_2 q_{xx} + i a_3 q_{xxx} + a_4 q_{xxxx} + i a_5 q_{xxxxx} + a_6 q_{xxxxxx} + (b_1 |q|^n + b_2 |q|^{2n} + b_3 |q|^{3n} + b_4 |q|^{4n} + b_5 |q|^{5n} + b_6 |q|^{6n}) q = i [\lambda (|q|^{2m} q)_x + \theta (|q|^{2m})_x q + \mu |q|^{2m} q_x]. \tag{1}$$

Here, q is the dependent variable standing for the complex-valued wave function which falls in the scope of nonlinear quantum optics. The coefficients a_j ($1 \leq j \leq 6$) that are real-valued represent dispersive effects and stem from inter-modal dispersion, chromatic dispersion, third-, fourth-, fifth- and sixth-order dispersion effects respectively. Next, b_j for $1 \leq j \leq 6$ are the six nonlinear forms of refractive index that was proposed by Kudryashov. The parameter n attributes to power-law effect. The first term in (1) refers to linear temporal evolution whose coefficient is $i = \sqrt{-1}$. From right side of the model, λ gives the effect of self-steepening while θ and μ represent nonlinear dispersion. Once again, the power-law parameter m accounts for maximum intensity.

This model has been very recently studied with the right side set to zero namely its unperturbed version [14]. In this case, bright, dark and singular solitons have been identified. However, (1) with right side set to zero will yield a bright 1-soliton solution for $n = 2$. Thus, the unperturbed equation (1) that yields bright 1-soliton solution is:

$$i q_t + i a_1 q_x + a_2 q_{xx} + i a_3 q_{xxx} + a_4 q_{xxxx} + i a_5 q_{xxxxx} + a_6 q_{xxxxxx} + (b_1 |q|^2 + b_2 |q|^4 + b_3 |q|^6 + b_4 |q|^8 + b_5 |q|^{10} + b_6 |q|^{12}) q = 0, \tag{2}$$

whose bright 1-soliton is as below [19]:

$$q(x, t) = A \operatorname{sech}^{\frac{1}{2}} [B(x - vt)] e^{i(-\kappa x + \omega t + \theta_0)}, \tag{3}$$

where A is the soliton amplitude, while B gives its inverse width and v stands for the soliton speed. Then, κ , ω and θ_0 respectively represent the frequency, wave number and phase constant.

The current work will study the perturbed version of Eq. (2) that is framed as:

$$i q_t + i a_1 q_x + a_2 q_{xx} + i a_3 q_{xxx} + a_4 q_{xxxx} + i a_5 q_{xxxxx} + a_6 q_{xxxxxx} + (b_1 |q|^2 + b_2 |q|^4 + b_3 |q|^6 + b_4 |q|^8 + b_5 |q|^{10} + b_6 |q|^{12}) q = i [\lambda (|q|^{2m} q)_x + \theta (|q|^{2m})_x q + \mu |q|^{2m} q_x]. \tag{4}$$

Eq. (4) will be integrated by the aid of SVP to locate its bright 1-soliton solution. This will yield an analytical solution for the governing model studied in (4) that is not exact.

Mathematical foundations

To commence, the starting presumptive hypothesis is [1-7,11-13]

$$q(x, t) = g(s) e^{i\phi(x,t)}, \tag{5}$$

where $g(s)$ stands for the amplitude component of the wave while $\phi(x, t)$ is its phase part. Here,

$$s = x - vt, \text{ and } \phi(x, t) = -\kappa x + \omega t + \theta_0.$$

Substitute (5) into (4). Then, imaginary part yields:

$$[v - a_1 + 2a_2\kappa + 3a_3\kappa^2 - 4a_4\kappa^3 + 5a_5\kappa^4 - 6a_6\kappa^5 + \{(2m + 1)\lambda + 2m\theta + \mu\} g^2] g' - (a_3 - 4a_4\kappa - 10a_5\kappa^2 + 20a_6\kappa^3) g''' - (a_5 - 6a_6\kappa) g^{(v)} = 0. \tag{6}$$

This implies the speed of the soliton is:

$$v = a_1 - 2a_2\kappa - 3a_3\kappa^2 + 4a_4\kappa^3 - 5a_5\kappa^4 + 6a_6\kappa^5. \tag{7}$$

The following parameter restrictions also come out of (6):

$$(2m + 1)\lambda + 2m\theta + \mu = 0, \tag{8}$$

$$a_3 - 4a_4\kappa - 10a_5\kappa^2 + 20a_6\kappa^3, \tag{9}$$

and

$$a_5 - 6a_6\kappa = 0. \tag{10}$$

The real part now implies

$$P_1 g - P_2 g'' + P_3 g^{(iv)} - a_6 g^{(vi)} + \kappa(\lambda + \mu) g^{2m+1} - b_1 g^3 - b_2 g^5 - b_3 g^7 - b_4 g^9 - b_5 g^{11} - b_6 g^{13} = 0, \tag{11}$$

with the notations

$$P_1 = \omega - a_1\kappa + a_2\kappa^2 + a_3\kappa^3 - a_4\kappa^4 - a_5\kappa^5 + a_6\kappa^6, \tag{12}$$

$$P_2 = a_2 + 3a_3\kappa - 6a_4\kappa^2 - 10a_5\kappa^3 + 15a_6\kappa^4, \tag{13}$$

and

$$P_3 = a_4 + 5a_5\kappa - 15a_6\kappa^2, \tag{14}$$

are implied. Multiplying (11) by g' and integrating leaves us with:

$$420P_1 g^2 - 420P_2 (g')^2 + 420P_3 (g'')^2 - 420a_6 (g''')^2 + \frac{420\kappa(\lambda + \mu)}{m + 1} g^{2m+2} - 210b_1 g^4 - 140b_2 g^6 - 105b_3 g^8 - 84b_4 g^{10} - 70b_5 g^{12} - 60b_6 g^{14} = K, \tag{15}$$

where K is the integration constant.

Implementation of SVP

The stationary integral is now defined as [1-7,11-13]:

$$J = \int_{-\infty}^{\infty} \left\{ 420P_1 g^2 - 420P_2 (g')^2 + 420P_3 (g'')^2 - 420a_6 (g''')^2 + \frac{420\kappa(\lambda + \mu)}{m + 1} g^{2m+2} - 210b_1 g^4 - 140b_2 g^6 - 105b_3 g^8 - 84b_4 g^{10} - 70b_5 g^{12} - 60b_6 g^{14} \right\} dx \tag{16}$$

From (3), we observe that

$$g(s) = A \operatorname{sech}^{\frac{1}{2}} (Bs). \tag{17}$$

Substituting (17) into (16) and performing the integration gives

$$J = \frac{5040\pi P_1 A^2}{B} - 630\pi P_2 A^2 B - 630\pi P_3 A^2 B^3 - 315\pi a_6 A^2 B^5 + \frac{10,080\kappa(\lambda + \mu) G A^{2m+2}}{m(m + 1) B} - \frac{5040b_1 A^4}{B} - \frac{840\pi b_2 A^6}{B} - \frac{1680b_3 A^8}{B} - \frac{378\pi b_4 A^{10}}{B} - \frac{896b_5 A^{12}}{B} - \frac{225\pi b_6 A^{14}}{B}, \tag{18}$$

where the notation

$$G = \frac{\Gamma\left(\frac{m+1}{2}\right) \Gamma\left(\frac{1}{2}\right)}{\Gamma\left(\frac{m}{2}\right)} \tag{19}$$

was adopted.

SVP points out that the bright 1-soliton solution to (4) is the same as the solution to (2) as given. However, its amplitude (A) and inverse width (B) will differ. In fact, they will be given by the coupled system of equations [1-7,11-13]:

$$\frac{\partial J}{\partial A} = 0, \tag{20}$$

and

$$\frac{\partial J}{\partial B} = 0, \tag{21}$$

For the reduced stationary integral as given by (18), the coupled system of Eqs. (20) and (21) formulates as:

$$\begin{aligned} &1680\pi P_1 - 210\pi P_2 B^2 + 210\pi P_3 B^4 - 105\pi a_6 B^6 + \frac{3360\kappa(\lambda + \mu)GA^{2m}}{m} \\ &- 3360b_1 A^2 - 560\pi b_2 A^4 - 2240b_3 A^6 - 630\pi b_4 A^8 \\ &- 1792b_5 A^{10} - 525\pi b_6 A^{12} = 0, \end{aligned} \tag{22}$$

and

$$\begin{aligned} &5040\pi P_1 + 630\pi P_2 B^2 + 1890\pi P_3 B^4 + 1575\pi a_6 B^6 \\ &+ \frac{10,080\kappa(\lambda + \mu)GA^{2m}}{m(m+1)} \\ &- 540b_1 A^2 - 840\pi b_2 A^4 - 1680b_3 A^6 - 378\pi b_4 A^8 \\ &- 896b_5 A^{10} - 225\pi b_6 A^{12} = 0, \end{aligned} \tag{23}$$

respectively. Upon uncoupling, one acquires the hexic polynomial equation for the inverse width B as:

$$\begin{aligned} &630\pi P_2 B^2 + 630\pi P_3 B^4 + 945\pi a_6 B^6 + \frac{540\kappa(\lambda + \mu)GA^{2m}}{m} \\ &+ 2520b_1 A^2 + 420\pi b_2 A^4 + 2520b_3 A^6 + 756\pi b_4 A^8 \\ &+ 2240b_5 A^{10} + 675\pi b_6 A^{12} = 0. \end{aligned} \tag{24}$$

Eq. (24) can be rewritten as the cubic polynomial equation:

$$au^3 + bu^2 + cu + d = 0, \tag{25}$$

where the following notations are utilized:

$$B^2 = u, \tag{26}$$

$$a = 945\pi a_6, \tag{27}$$

$$b = 630\pi P_3, \tag{28}$$

$$c = 630\pi P_2, \tag{29}$$

$$\begin{aligned} d = &\frac{540\kappa(\lambda + \mu)GA^{2m}}{m} + 2520b_1 A^2 + 420\pi b_2 A^4 + 2520b_3 A^6 \\ &+ 756\pi b_4 A^8 + 2240b_5 A^{10} + 675\pi b_6 A^{12}. \end{aligned} \tag{30}$$

While (19) implicates that $m > 0$, one concludes from (30),

$$m \neq 0. \tag{31}$$

By virtue of Cardano's method, applied to (25) and subsequently from (26), one recovers the width of the soliton in terms of its amplitude as

$$\begin{aligned} B = &\left[\left\{ \left(-\frac{b^3}{27a^3} + \frac{bc}{6a^2} - \frac{d}{2a} \right) \right. \right. \\ &- \sqrt{\left(-\frac{b^3}{27a^3} + \frac{bc}{6a^2} - \frac{d}{2a} \right)^2 + \left(\frac{c}{3a} - \frac{b^2}{9a^2} \right)^3} \Bigg]^{\frac{1}{3}} \\ &+ \left\{ \left(-\frac{b^3}{27a^3} + \frac{bc}{6a^2} - \frac{d}{2a} \right) \right. \\ &+ \sqrt{\left(-\frac{b^3}{27a^3} + \frac{bc}{6a^2} - \frac{d}{2a} \right)^2 + \left(\frac{c}{3a} - \frac{b^2}{9a^2} \right)^3} \Bigg]^{\frac{1}{3}} - \frac{b}{3a} \Bigg]^{\frac{1}{2}}. \end{aligned} \tag{32}$$

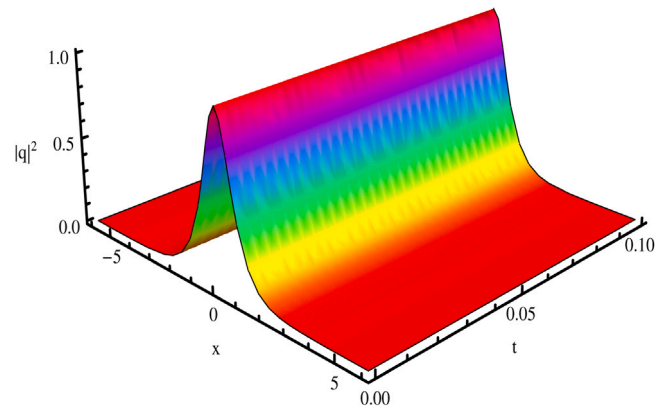


Fig. 1. $|q|^2$ soliton.

The constraints for this solution to exist are:

$$a_6 \neq 0, \tag{33}$$

along with the discriminant

$$\left(-\frac{b^3}{27a^3} + \frac{bc}{6a^2} - \frac{d}{2a} \right)^2 + \left(\frac{c}{3a} - \frac{b^2}{9a^2} \right)^3 > 0, \tag{34}$$

as well as

$$\begin{aligned} &\left\{ \left(-\frac{b^3}{27a^3} + \frac{bc}{6a^2} - \frac{d}{2a} \right) - \sqrt{\left(-\frac{b^3}{27a^3} + \frac{bc}{6a^2} - \frac{d}{2a} \right)^2 + \left(\frac{c}{3a} - \frac{b^2}{9a^2} \right)^3} \right\}^{\frac{1}{3}} \\ &+ \left\{ \left(-\frac{b^3}{27a^3} + \frac{bc}{6a^2} - \frac{d}{2a} \right) + \sqrt{\left(-\frac{b^3}{27a^3} + \frac{bc}{6a^2} - \frac{d}{2a} \right)^2 + \left(\frac{c}{3a} - \frac{b^2}{9a^2} \right)^3} \right\}^{\frac{1}{3}} \\ &> \frac{b}{3a}. \end{aligned} \tag{35}$$

The solution to (4) is given by (3), where the soliton width B is related to the amplitude A as in (32) with all of the notations and constraints as discussed in the work.

Fig. 1 demonstrates the numerical simulation of a bright 1-soliton as given by (3) for perturbed Kudryashov's model. The system parameter values are: $m = 1$, $a_1 = 1$, $a_2 = 0.1$, $a_4 = 2$, $a_6 = -0.3$, $b_j = 0.1$ ($j = 1, 2, \dots, 6$), $\omega = 1$, $\kappa = \lambda = \mu = 0.5$, $A = 1$.

Conclusions

This paper is about SVP applied successfully to perturbed Kudryashov's equation that contains six nonlinear forms of SPM. In this context, bright HD 1-soliton solution is revealed. This analytical solution is in its closed form although not an exact solution. A version of the variational principle, commonly referred to as SVP, has made this retrieval possible. A numerical simulation is also presented. The perturbation terms are with maximum intensity. No known integration scheme can possibly identify a closed form solution. Thus, SVP is to the rescue. One other shortcoming, however, of the scheme is that it fails to retrieve dark or singular soliton solution. It is the stationary integral that is rendered divergent and thus the current approach meets its limitation. Later, this approach will be implemented with other models as introduced by Kudryashov that will lead to its bright 1-soliton solution. They are surely going to be visible somewhere. The readers are suggested to stay tuned!

CRedit authorship contribution statement

Anjan Biswas: Supervision, Investigation; Writing – original draft, Methodology, Software. **Mehmet Ekici:** Supervision, Investigation,

Writing – original draft. **Anelia Dakova:** Validation, Investigation. **Salam Khan:** Supervision, Validation. **Seithuti P. Moshokoa:** Supervision, Validation. **Hashim M. Alshehri:** Writing – review & editing. **Milivoj R. Belic:** Writing – review & editing, Funding acquisition.

Declaration of competing interest

The authors declare that they have no known competing financial interests or personal relationships that could have appeared to influence the work reported in this paper.

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